

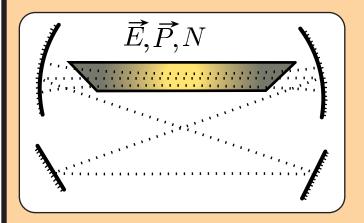
Mechanical analog of a laser close to the 1:1 resonance

Marcel G. Clerc Departamento de Fisica Universidad de Chile

Outline

- Introduction to the laser Physics.
- Laser Instability.
- Hamiltonian bifurcation and quasi-reversible instability.
- Examples of 1:1 Resonances
- Mechanical analog of the laser
- Latent Bifurcation.
- Conclusion.

Semiclassical Laser Model



 $\begin{aligned} \frac{\partial^2 E}{\partial t^2} &= \frac{\partial^2 E}{\partial x^2} - \frac{\partial^2 P}{\partial t^2} - \kappa \frac{\partial E}{\partial t}, \\ \frac{\partial^2 P}{\partial t^2} &= -\gamma_{\perp} \frac{\partial P}{\partial t} - (\gamma_{\perp}^2 + \Omega^2) P - \mu^2 N E, \\ \frac{\partial N}{\partial t} &= -\gamma_{\parallel} (N - N_0) + E \left(\frac{\partial P}{\partial t} + \gamma_{\perp} P\right), \end{aligned}$

where the terms proportional to κ, γ_{\perp} and γ_{\parallel} are dissipatives. The nonlasing solution (equilibrium) is

$$E = P = 0$$
 and $N = D_0$

Non dissipative semiclassical model

$$\frac{\partial^2 E}{\partial t^2} = \frac{\partial^2 E}{\partial x^2} - \left(\frac{\mu}{\Omega}\right)^2 \frac{\partial^2 P}{\partial t^2},\\ \frac{\partial P}{\partial t} = G, \quad \frac{\partial G}{\partial t} = -P - NE,\\ \frac{\partial N}{\partial t} = EG.$$

where $P \rightarrow \mu/\Omega P$, $E \rightarrow \Omega/\mu E$, $t \rightarrow t/\Omega$, $x \rightarrow x/\Omega$ This model has Hamiltonian density

$$\mathcal{H} = \frac{1}{2} \left(D - \left(\frac{\mu}{\Omega}\right)^2 P \right)^2 + \frac{\left(\partial_x A\right)^2}{2} + \left(\frac{\mu}{\Omega}\right)^2 N,$$

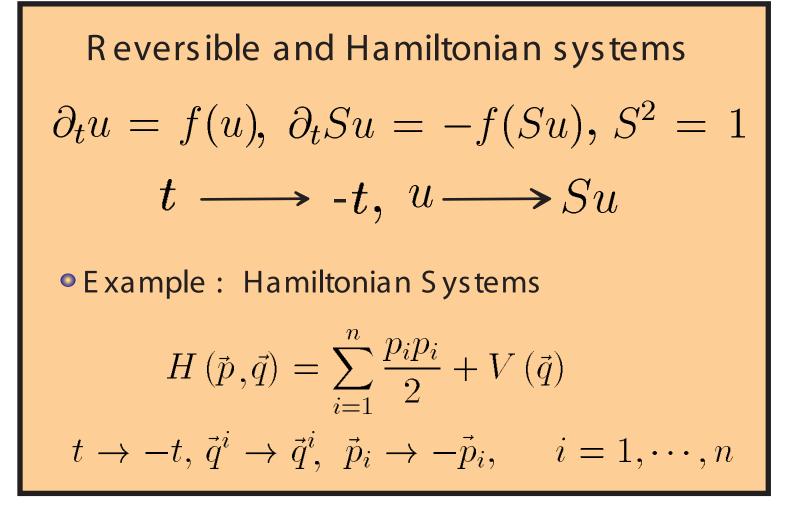
with $E = \partial_t A = D - (\mu/\Omega)^2 P$.

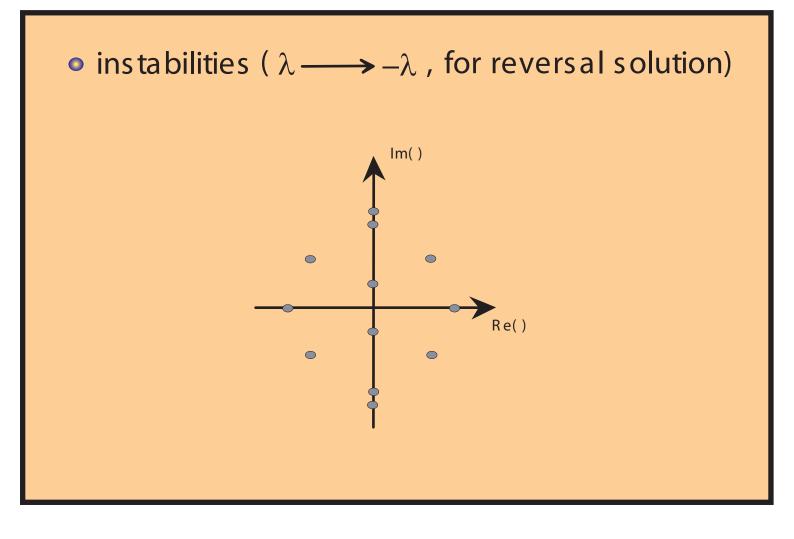
and the Poisson-bracket $\{F, K\} = \int dx \left\{ \frac{\partial F}{\partial A} \frac{\partial K}{\partial D} - \frac{\partial F}{\partial D} \frac{\partial K}{\partial A} - \left(\frac{\mu}{\Omega}\right)^2 \vec{m} \cdot \left(\vec{\nabla}_m F \times \vec{\nabla}_m K\right) \right\},$ where $\vec{m} = (N, P, G)$

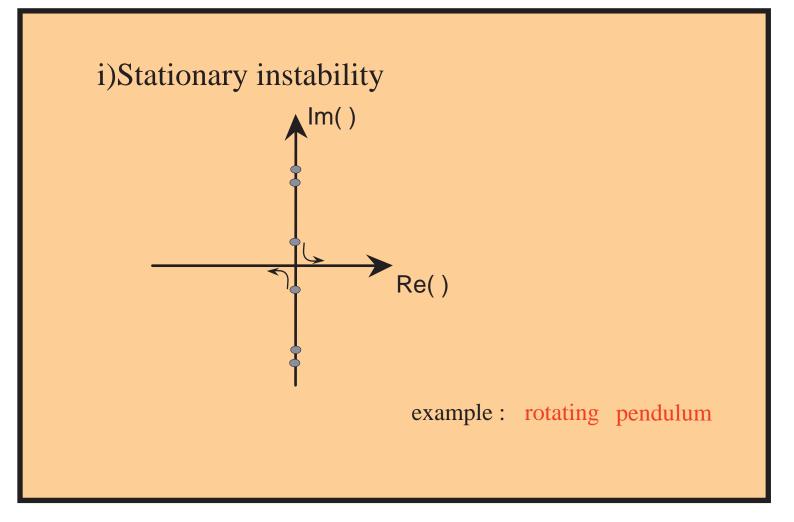
Energie-Casimir Method

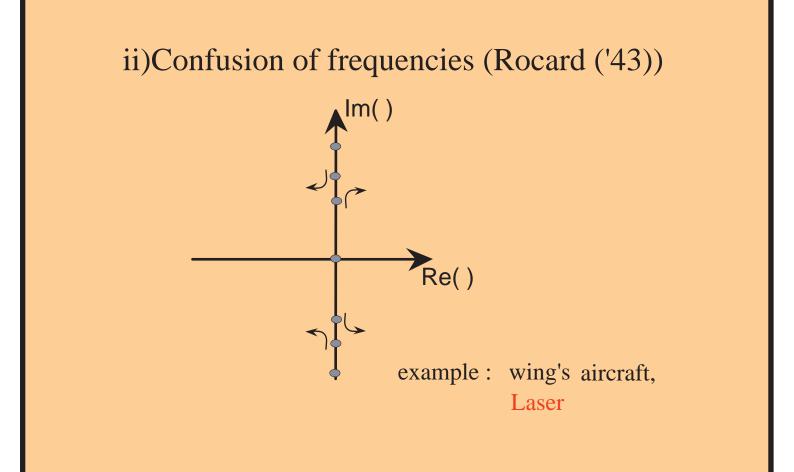
• $\{\Phi(N^2 + P^2 + G^2), F\} = 0$ • The effective energie

$$H_{C} = \int \left\{ \frac{1}{2} \left(D - \left(\frac{\mu}{\Omega}\right)^{2} P \right)^{2} + \frac{(\partial_{x} A)^{2}}{2} + \left(\frac{\mu}{\Omega}\right)^{2} N + \left(\frac{\mu}{\Omega}\right)^{2} \frac{(N^{2} + P^{2} + G^{2})}{2D_{o}} + \alpha^{2} (N^{2} + P^{2} + G^{2})^{2} \right\},$$

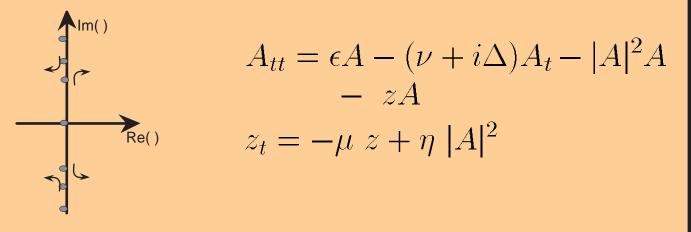








 In the quasi-reversible confusion of frequencies, the asymptotic normal form



 J. Gibbon et al ('80): the dispersive instability with small dissipation.

Introducing the change of variables

$$P = \kappa E + \partial_t E, \quad E = e^{-i\frac{\Delta\kappa}{\gamma+\kappa}t}\frac{A}{\sqrt{g}}, \quad N = \frac{z - D_o}{g} + |E|^2,$$

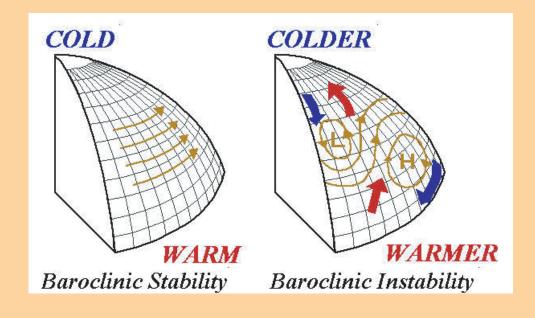
the equations read (Maxwell-Bloch or complex Lorenz eqs.)

$$\begin{split} \partial_t E &= -\kappa E + P \\ \partial_t P &= -(\gamma_\perp + i\Delta)P - gNE \\ \partial_t N &= -\gamma_{||}(N - N_0) + (E\bar{P} + \bar{E}P) \end{split}$$

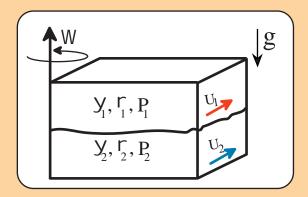
Physical examples

Baroclinic instability

This instability gives rise waves motion due to vertical shear of the basic current in the presence of Coriolis and buoyancy forces



Quasi-Geostrophic Two-layer Model



two layers of immiscible, incompressible, homogeneous fluid of slightly different densities (r > r). The dimensionless quasi-geostrophic vorticity equations are

$$\begin{aligned} \left[\partial_t + \psi_{1,x}\partial_y - \psi_{1,y}\partial_x\right] \left[\vec{\nabla}^2\psi_1 + F\left(\psi_2 - \psi_1\right) + \beta y\right] &= -r\vec{\nabla}^2\psi_1, \\ \left[\partial_t + \psi_{2,x}\partial_y - \psi_{2,y}\partial_x\right] \left[\vec{\nabla}^2\psi_2 + F\left(\psi_1 - \psi_2\right) + \beta y\right] &= -r\vec{\nabla}^2\psi_2, \end{aligned}$$

• This system is Lagrangean

$$\mathcal{L} = \int \frac{1}{2} \left(\vec{\nabla} \psi_1 \right)^2 + \frac{1}{2} \left(\vec{\nabla} \psi_1 \right)^2 + \frac{F}{2} \left(\psi_1 - \psi_2 \right)^2 - \beta y \left(\psi_1 + \psi_2 \right) dt dx dy,$$

constrained to Euler-Poincaré variations

$$\delta\psi_1 = \frac{\partial}{\partial t}\delta\phi_1 + \hat{z}\cdot\left(\vec{\nabla}\psi_1\times\vec{\nabla}\delta\phi_1\right),\\ \delta\psi_2 = \frac{\partial}{\partial t}\delta\phi_2 + \hat{z}\cdot\left(\vec{\nabla}\psi_2\times\vec{\nabla}\delta\phi_2\right).$$

Where the stream functions depent of the horizontal coordinate and time. The respective Hamiltonian is

$$\mathcal{H} = \int \frac{1}{2} \left(\vec{\nabla} \psi_1 \right)^2 + \frac{1}{2} \left(\vec{\nabla} \psi_1 \right)^2 + \frac{F}{2} \left(\psi_1^2 + \psi_2^2 \right) dx dy,$$

If one considers perturbations of the geostrophic basic flow of the form

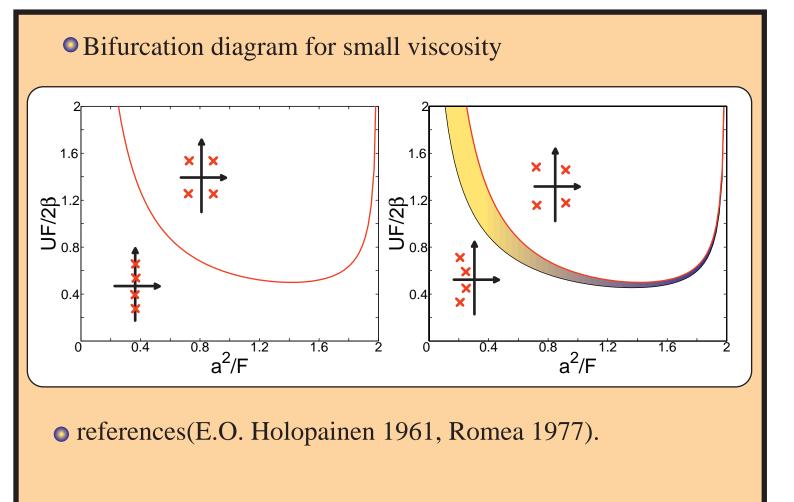
$$\psi_1 = -U_1 y + \operatorname{Re} A e^{i\alpha(x-ct)} \sin(m\pi y),$$

$$\psi_2 = -U_2 y + \operatorname{Re} \gamma A e^{i\alpha(x-ct)} \sin(m\pi y),$$

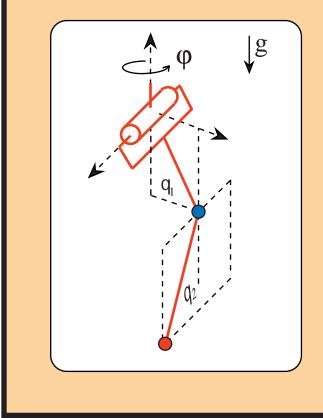
then dispertion relation is

$$c = \frac{U_1 + U_2}{2} - \frac{a^2 + F}{a^2 + 2F} \left[\frac{\beta + i\alpha r}{\alpha^2} \right]$$
$$\pm \frac{\left[(\Delta U)^2 a^4 \left(a^4 - 4F^2 \right) + 4F^2 \left(\beta + ira^2 \alpha^{-1} \right) \right]^{1/2}}{2a^2 \left(a^2 + 2F \right)}$$

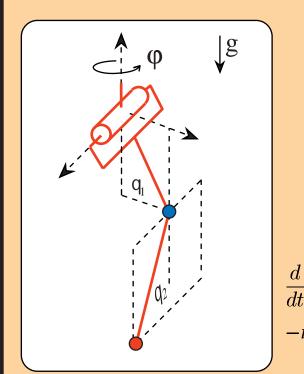
where $a^2 = \alpha^2 + m^2 \pi^2$ and $\Delta U \equiv U_1 - U_2$.



Mechanical Laser



Two coupled spherical pendula in a gravitational field, with a support, which can rotate around the vertical axis. The lower pendulum is constrained to move in a plane that is orthogonal to the plane of the upper pendulum.

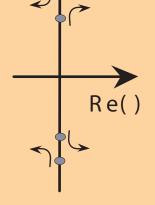


$$\begin{split} \ddot{\theta}_{1} &= -\sigma^{2} \sin \theta_{1} \sin \theta_{2} \ddot{\theta}_{2} - \sigma^{2} \sin \theta_{1} \cos \theta_{2} \dot{\theta}_{2}^{2} \\ &- 2\sigma^{2} \cos \theta_{1} \cos \theta_{2} \dot{\varphi} \dot{\theta}_{2} + \sin \theta_{1} \cos \theta_{1} \dot{\varphi}^{2} \\ &- \sigma^{2} \cos \theta_{1} \sin \theta_{2} \ddot{\varphi} - \frac{g}{l} \sin \theta_{1} - \nu_{1} \dot{\theta}_{1}, \\ \ddot{\theta}_{2} &= -\sin \theta_{1} \sin \theta_{2} \ddot{\theta}_{1} - \cos \theta_{1} \sin \theta_{2} \dot{\theta}_{1}^{2} \\ &+ 2\cos \theta_{1} \cos \theta_{2} \dot{\varphi} \dot{\theta}_{1} + \sin \theta_{2} \cos \theta_{2} \dot{\varphi}^{2} \\ &+ \sin \theta_{1} \cos \theta_{2} \ddot{\varphi} - \frac{g}{l} \sin \theta_{2} - \nu_{2} \dot{\theta}_{2}, \\ \ddot{e}_{1} \left\{ \begin{array}{c} (\sin^{2} \theta_{1} + \sigma^{2} \sin^{2} \theta_{2}) \dot{\varphi} + \sigma^{2} \cos \theta_{1} \sin \theta_{2} \dot{\theta}_{1} \\ &- \sigma^{2} \sin \theta_{1} \cos \theta_{2} \dot{\theta}_{2} + I \dot{\varphi} \end{array} \right\} = \\ \nu_{\varphi} \left(\dot{\varphi} - \Omega \right) - \mu_{1} \sin^{2} \theta_{1} \dot{\varphi} - \mu_{2} \left(\sin^{2} \theta_{1} + \sin^{2} \theta_{2} \right) \dot{\varphi}. \end{split}$$

where $l_1 = l_2 = l$ and $\sigma = \sqrt{m_2 / (m_1 + m_2)}$



$$\theta_1 = \theta_2 = 0, \, \varphi_t = \Omega_0$$



Im()

exhibits a 1:1 resonance when

$$\Omega_0 = \Omega_c = \sqrt{\frac{g}{l} \frac{(m_1 + m_2)}{m_1}}$$

with $\omega_c = \pm \sqrt{g m_2 / l m_1}$ frequencies. The system is described by

$$A_{tt} = \frac{2g\left(\Omega - \Omega_c\right)}{l\Omega} A + i\left(2\sigma\left(\Omega - \Omega_c\right)\right) A_t$$

• Quasi-reversible instability

$$\partial_{tt}A = (\varepsilon - Z)A - (\mu - i\delta)\partial_t A - \alpha |A|^2 A$$

$$\partial_t Z = \nu Z + \eta |A|^2.$$

where

$$\varepsilon = 2\frac{g}{l}\frac{(\Omega - \Omega_c)}{\Omega_c}, \ \alpha = \frac{g}{4l}\left(\frac{\sigma^4 - 2\sigma^3 - 2\sigma^2 + 3}{1 - \sigma^2}\right),$$

$$\delta = 2\sigma\left(\Omega - \Omega_c\right), \ \nu = \frac{\nu_{\varphi}}{I}, \ \mu = \frac{1}{2l^2}\left(\nu_1 + \nu_2\right),$$

$$\eta = \frac{1}{I}\left[\mu_1 + \left(1 + \frac{1}{\sigma^2}\right)\mu_2 - 2\frac{\nu_{\varphi}}{\sigma^2 I}\left(\frac{\Omega}{\Omega_c} - 1\right)\right];$$

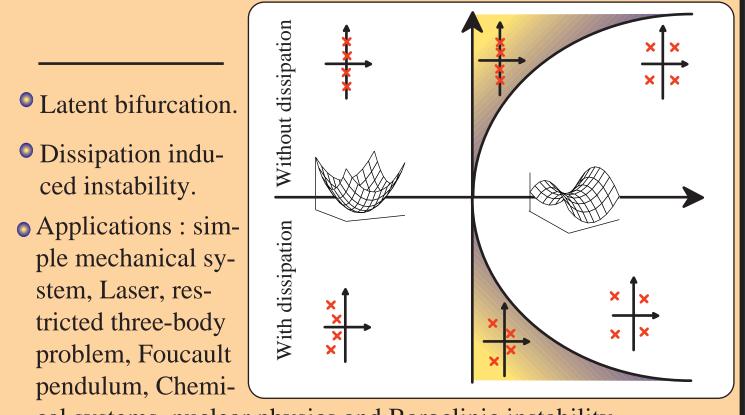
FOUCAULT PENDULUM





• As consequence of Earth rotation, the vertical solution exhibits a latent bifurcation!.

$$\begin{split} \ddot{x} &= -\left(\frac{g}{l} - \Omega^2\right)x + 2\Omega \dot{y} \\ \ddot{y} &= -\left(\frac{g}{l} - \Omega^2\right)y - 2\Omega \dot{x} \end{split}$$



cal systems, nuclear physics and Baroclinic instability.