

Latent Bifurcation



"Latent Bifurcation, Dissipation induce instability and applications".

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Outline

- Examples.
- Definition of linearly and spectrally instability.
- Dynamics around the 1:1 resonance.
- Latent bifurcation.
- Dissipation induced instability.
- Applications.

Mechanical Laser



Two coupled spherical pendula in a gravitational field, with a support, which can rotate around the vertical axis. The lower pendulum is constrained to move in a plane that is orthogonal to the plane of the upper pendulum.



$$\begin{split} \ddot{\theta}_{1} &= -\sigma^{2} \sin \theta_{1} \sin \theta_{2} \ddot{\theta}_{2} - \sigma^{2} \sin \theta_{1} \cos \theta_{2} \dot{\theta}_{2}^{2} \\ &- 2\sigma^{2} \cos \theta_{1} \cos \theta_{2} \dot{\varphi} \dot{\theta}_{2} + \sin \theta_{1} \cos \theta_{1} \dot{\varphi}^{2} \\ &- \sigma^{2} \cos \theta_{1} \sin \theta_{2} \ddot{\varphi} - \frac{g}{l} \sin \theta_{1} - \nu_{1} \dot{\theta}_{1}, \\ \ddot{\theta}_{2} &= -\sin \theta_{1} \sin \theta_{2} \ddot{\theta}_{1} - \cos \theta_{1} \sin \theta_{2} \dot{\theta}_{1}^{2} \\ &+ 2\cos \theta_{1} \cos \theta_{2} \dot{\varphi} \dot{\theta}_{1} + \sin \theta_{2} \cos \theta_{2} \dot{\varphi}^{2} \\ &+ \sin \theta_{1} \cos \theta_{2} \ddot{\varphi} - \frac{g}{l} \sin \theta_{2} - \nu_{2} \dot{\theta}_{2}, \\ \ddot{e}_{1} \left\{ \begin{array}{c} (\sin^{2} \theta_{1} + \sigma^{2} \sin^{2} \theta_{2}) \dot{\varphi} + \sigma^{2} \cos \theta_{1} \sin \theta_{2} \dot{\theta}_{1} \\ &- \sigma^{2} \sin \theta_{1} \cos \theta_{2} \dot{\theta}_{2} + I \dot{\varphi} \end{array} \right\} = \\ \nu_{\varphi} \left(\dot{\varphi} - \Omega \right) - \mu_{1} \sin^{2} \theta_{1} \dot{\varphi} - \mu_{2} \left(\sin^{2} \theta_{1} + \sin^{2} \theta_{2} \right) \dot{\varphi}. \end{split}$$

where $l_1 = l_2 = l$ and $\sigma = \sqrt{m_2 / (m_1 + m_2)}$



Semiclassical Laser Model



$$\frac{\partial^2 E}{\partial t^2} = \frac{\partial^2 E}{\partial x^2} - \frac{\partial^2 P}{\partial t^2} - \kappa \frac{\partial E}{\partial t},$$

$$\frac{\partial^2 P}{\partial t^2} = -\gamma_{\perp} \frac{\partial P}{\partial t} - (\gamma_{\perp}^2 + \Omega^2) P - \mu^2 N E,$$

$$\frac{\partial N}{\partial t} = -\gamma_{\parallel} (N - N_0) + E \left(\frac{\partial P}{\partial t} + \gamma_{\perp} P\right),$$

where the terms proportional to κ, γ_{\perp} and γ_{\parallel} are dissipatives. The nonlasing solution (equilibrium) is

$$E = P = 0$$
 and $N = D_0$

Stability of Equilibria

When all solutions of linearized equation at equilibrium are stable (Lyapunov), the equilibrium is said to be linearly stable or linearized stable.







The spectral instability can predict instability, but not stability!

1:1 Resonance



Nearby this instability the system is decribed by the normal form

$$\partial_t A = i\Omega A + B$$

$$\partial_t B = i\Omega B + f(|A|^2, i(AB^* - BA^*), \{\lambda\}) A + ig(|A|^2, i(AB^* - BA^*), \{\lambda\}) B,$$

where f and g are real functions, and {|} is a set of parameters.

The linear evolution around the zero solution $(\mathcal{A} = \mathcal{B} = 0)$

$$\partial_t \mathcal{A} = i\Omega \mathcal{A} + \mathcal{B},$$

$$\partial_t \mathcal{B} = i\left(\Omega + \delta\right) \mathcal{B} + \varepsilon \mathcal{A},$$

introducing the rotating variables $\mathcal{A} = Ae^{i\Omega t} = (x + iy)e^{i\Omega t}$ and $\mathcal{B} = \partial_t Ae^{i\Omega t} = (\partial_t x + i\partial_t y)e^{i\Omega t}$

$$\partial_{tt} x = \varepsilon x - \delta \partial_t y,$$

$$\partial_{tt} y = \varepsilon y + \delta \partial_t x.$$

• Gyroscopic system.

$$\partial_{tt}A = \varepsilon A + i\delta\partial_t A$$

The gyscopic system is hamiltonian

$$H = \partial_t A \partial_t A^* - \varepsilon |A|^2$$

with the Poisson-bracket

$$\{F,G\} = \frac{\partial F}{\partial A} \frac{\partial G}{\partial A_t^*} - \frac{\partial G}{\partial A} \frac{\partial F}{\partial A_t^*} + i\delta \frac{\partial F}{\partial A_t} \frac{\partial G}{\partial A_t} + \text{c.c.}$$

The eigenvalues of this equilibrium are

$$\sigma = \pm 1/2\sqrt{4\varepsilon - 2\delta^2 \pm 2\delta\sqrt{\delta^2 - 4\varepsilon}}$$

Bifurcation Diagram





Latent Bifurcation



The instability is present but non perceptible with the spectrum, and requieres a large time in come into view.

There is not resonance between the frequencies.

Dissipation induced instability

The equations under the presence of small dissipative terms read (m<<1) $\partial_{tt}x = \varepsilon x - \delta \partial_t y - \mu \partial_t x,$ $\partial_{tt}y = \varepsilon y + \delta \partial_t x - \mu \partial_t y,$ • Bifurcation diagram $d = \varepsilon x - \delta \partial_t y - \mu \partial_t x,$ $\partial_{tt}y = \varepsilon y + \delta \partial_t x - \mu \partial_t y,$

Observations



- The unperturbed system is marginal, the eigenvalues with larger frequency move to the left of the imaginary axis and these are the furthest from this axis.
- The destabilizing effects through positive or negative total dissipative perturbation was know a long time ago (Lord Kelvin, 1897).
- The asymptotic normal form of 1:1 resonance

$$\partial_{tt}A = \varepsilon A - (\mu - i\delta)\,\partial_t A - \alpha |A|^2 A,$$

Mechanical Laser



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where $l_1 = l_2 = l$ and $\sigma = \sqrt{m_2 / (m_1 + m_2)}$



Semiclassical Laser Model



$$\frac{\partial^2 E}{\partial t^2} = \frac{\partial^2 E}{\partial x^2} - \frac{\partial^2 P}{\partial t^2} - \kappa \frac{\partial E}{\partial t},$$

$$\frac{\partial^2 P}{\partial t^2} = -\gamma_{\perp} \frac{\partial P}{\partial t} - (\gamma_{\perp}^2 + \Omega^2) P - \mu^2 N E,$$

$$\frac{\partial N}{\partial t} = -\gamma_{\parallel} (N - N_0) + E \left(\frac{\partial P}{\partial t} + \gamma_{\perp} P\right),$$

where the terms proportional to κ_{γ} , γ_{\perp} and γ_{\parallel} are dissipatives. The nonlasing solution (equilibrium) is

$$E = P = 0$$
 and $N = D_0$

Non dissipative semiclassical model

$$\frac{\partial^2 E}{\partial t^2} = \frac{\partial^2 E}{\partial x^2} - \left(\frac{\mu}{\Omega}\right)^2 \frac{\partial^2 P}{\partial t^2},\\ \frac{\partial P}{\partial t} = G, \quad \frac{\partial G}{\partial t} = -P - NE,\\ \frac{\partial N}{\partial t} = EG.$$

where $P \rightarrow \mu/\Omega P$, $E \rightarrow \Omega/\mu E$, $t \rightarrow t/\Omega$, $x \rightarrow x/\Omega$ This model has Hamiltonian density

$$\mathcal{H} = \frac{1}{2} \left(D - \left(\frac{\mu}{\Omega}\right)^2 P \right)^2 + \frac{\left(\partial_x A\right)^2}{2} + \left(\frac{\mu}{\Omega}\right)^2 N,$$

with $E = \partial_t A = D - (\mu/\Omega)^2 P$.

and the Poisson-bracket $\{F, K\} = \int dx \left\{ \frac{\partial F}{\partial A} \frac{\partial K}{\partial D} - \frac{\partial F}{\partial D} \frac{\partial K}{\partial A} - \left(\frac{\mu}{\Omega}\right)^2 \vec{m} \cdot \left(\vec{\nabla}_m F \times \vec{\nabla}_m K\right) \right\},$ where $\vec{m} = (N, P, G)$

Energie-Casimir Method

• $\{\Phi(N^2 + P^2 + G^2), F\} = 0$ • The effective energie

$$H_{C} = \int \left\{ \frac{1}{2} \left(D - \left(\frac{\mu}{\Omega}\right)^{2} P \right)^{2} + \frac{(\partial_{x} A)^{2}}{2} + \left(\frac{\mu}{\Omega}\right)^{2} N + \left(\frac{\mu}{\Omega}\right)^{2} \frac{(N^{2} + P^{2} + G^{2})}{2D_{o}} + \alpha^{2} (N^{2} + P^{2} + G^{2})^{2} \right\},$$



Thus, a slighty pumping optical cavity is unstable ($D_0>0$) when one take into account the dissipative effects and the nonlassing solution exhibits a latent bifurcation for D_0 equal to zero.

Baroclinic instability

This instability gives rise waves motion due to vertical shear of the basic current in the presence of Coriolis and buoyancy forces



Quasi-Geostrophic Two-layer Model



two layers of immiscible, incompressible, homogeneous fluid of slightly different densities (r > r). The dimensionless quasi-geostrophic vorticity equations are

$$\begin{aligned} \left[\partial_t + \psi_{1,x}\partial_y - \psi_{1,y}\partial_x\right] \left[\vec{\nabla}^2\psi_1 + F\left(\psi_2 - \psi_1\right) + \beta y\right] &= -r\vec{\nabla}^2\psi_1, \\ \left[\partial_t + \psi_{2,x}\partial_y - \psi_{2,y}\partial_x\right] \left[\vec{\nabla}^2\psi_2 + F\left(\psi_1 - \psi_2\right) + \beta y\right] &= -r\vec{\nabla}^2\psi_2, \end{aligned}$$

• This system is Lagrangean

$$\mathcal{L} = \int \frac{1}{2} \left(\vec{\nabla} \psi_1 \right)^2 + \frac{1}{2} \left(\vec{\nabla} \psi_1 \right)^2 + \frac{F}{2} \left(\psi_1 - \psi_2 \right)^2 - \beta y \left(\psi_1 + \psi_2 \right) dt dx dy,$$

constrained to Euler-Poincaré variations

$$\delta\psi_1 = \frac{\partial}{\partial t}\delta\phi_1 + \hat{z}\cdot\left(\vec{\nabla}\psi_1\times\vec{\nabla}\delta\phi_1\right),\\ \delta\psi_2 = \frac{\partial}{\partial t}\delta\phi_2 + \hat{z}\cdot\left(\vec{\nabla}\psi_2\times\vec{\nabla}\delta\phi_2\right).$$

Where the stream functions depent of the horizontal coordinate and time. The respective Hamiltonian is

$$\mathcal{H} = \int \frac{1}{2} \left(\vec{\nabla} \psi_1 \right)^2 + \frac{1}{2} \left(\vec{\nabla} \psi_1 \right)^2 + \frac{F}{2} \left(\psi_1^2 + \psi_2^2 \right) dx dy,$$

If one considers perturbations of the geostrophic basic flow of the form

$$\psi_1 = -U_1 y + \operatorname{Re} A e^{i\alpha(x-ct)} \sin(m\pi y),$$

$$\psi_2 = -U_2 y + \operatorname{Re} \gamma A e^{i\alpha(x-ct)} \sin(m\pi y),$$

then dispertion relation is

$$c = \frac{U_1 + U_2}{2} - \frac{a^2 + F}{a^2 + 2F} \left[\frac{\beta + i\alpha r}{\alpha^2}\right]$$
$$\pm \frac{\left[(\Delta U)^2 a^4 \left(a^4 - 4F^2\right) + 4F^2 \left(\beta + ira^2 \alpha^{-1}\right)\right]^{1/2}}{2a^2 \left(a^2 + 2F\right)}$$

where $a^2 = \alpha^2 + m^2 \pi^2$ and $\Delta U \equiv U_1 - U_2$.





Using synodic coordinates, that is, we consider the rotating reference frame with the same frequency as primaries, one finds that the evolution of the test particle is described by

$$egin{aligned} \ddot{x}-2\dot{y}&=&rac{\partial\Omega}{\partial x},\ \ddot{y}+2\dot{x}&=&rac{\partial\Omega}{\partial y}, \end{aligned}$$

where $\Omega(x,y) = \frac{x^2 + y^2}{2} + \frac{1-\mu}{r_1} + \frac{\mu}{r_2},$

$$r_1 = \sqrt{(x+\mu)^2 + y^2}$$
, and $r_2 = \sqrt{(x-1+\mu)^2 + y^2}$



• Equilateral libration points are marginal stable!, when m is small.

LAGRANGE POINTS

 Librations points in the planar circular restricted three-body problem. The stability of these points is classical problem in Celestial Mechanics.

$$x = \frac{1}{2} - \mu, \quad y = \pm \sqrt{\frac{3}{2}},$$

Dissipation induced instabilities \mathbf{r}

Inertial drag force

$$\ddot{x} - 2\dot{y} = \frac{\partial\Omega}{\partial x} - \nu \left(\dot{x} - y\right),$$

$$\ddot{y} + 2\dot{x} = \frac{\partial\Omega}{\partial y} - \nu \left(\dot{y} + x\right).$$

• Equilateral libration points are stable! (Murray 1994).

Dissipation induced instabilities



$$\ddot{x} - 2\dot{y} = rac{\partial\Omega}{\partial x} - \nu\dot{x},$$

 $\ddot{y} + 2\dot{x} = rac{\partial\Omega}{\partial y} - \nu\dot{y}.$

Nebular drag force

• Equilateral libration points are unstable!



FOUCAULT PENDULUM





• As consequence of Earth rotation, the vertical solution exhibits a latent bifurcation!.

$$\begin{split} \ddot{x} &= -\left(\frac{g}{l} - \Omega^2\right)x + 2\Omega \dot{y} \\ \ddot{y} &= -\left(\frac{g}{l} - \Omega^2\right)y - 2\Omega \dot{x} \end{split}$$

Conclusion

- Latent bifurcation.
- Dissipation induced instability.
- Applications : simple mechanical system, Laser, restricted three-body problem, Foucault pendulum, Chemi-



cal systems, nuclear physics and Baroclinic instability.